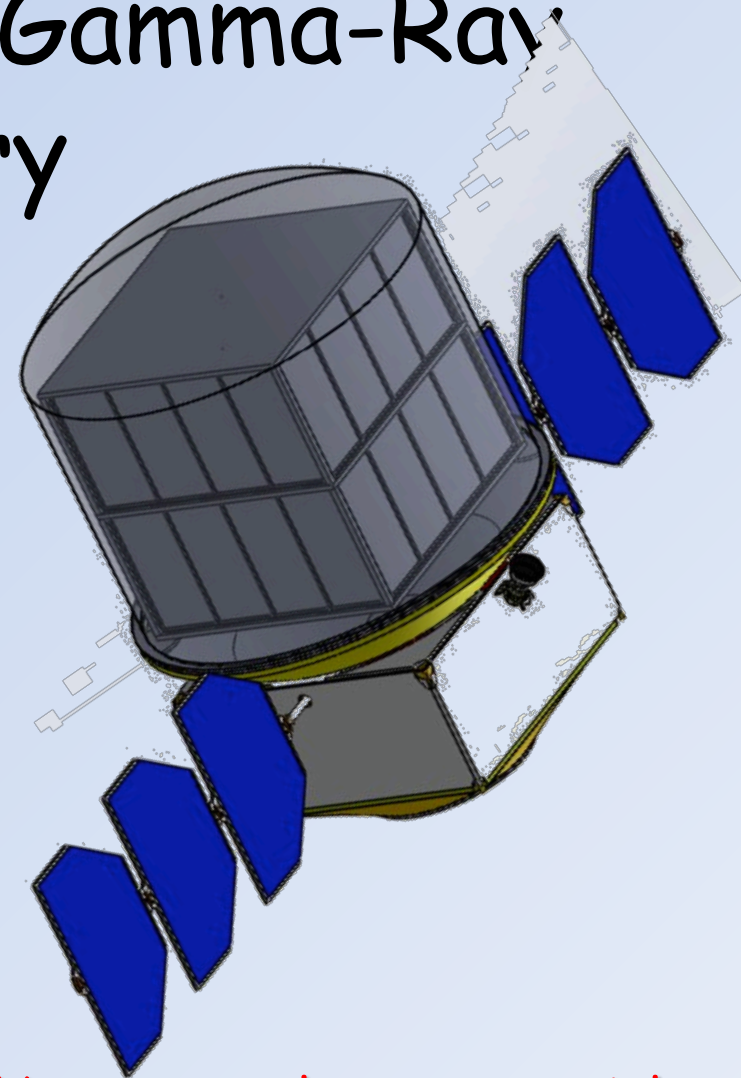


Adavanced Energetic Pair Telescope for Medium-Energy Gamma-Ray Polarimetry

Stanley D. Hunter
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Francesco Longo

Future Space-based Gamma-ray Obs.
GSFC, Feb. 6, 2015



Hunter, et al., *Astroparticle
Physics*, 59, 18-28 (2014)

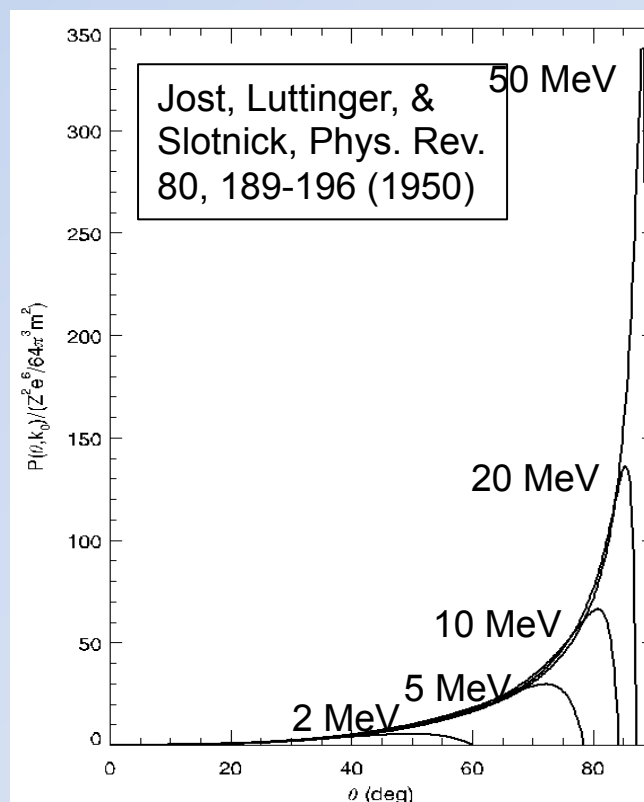
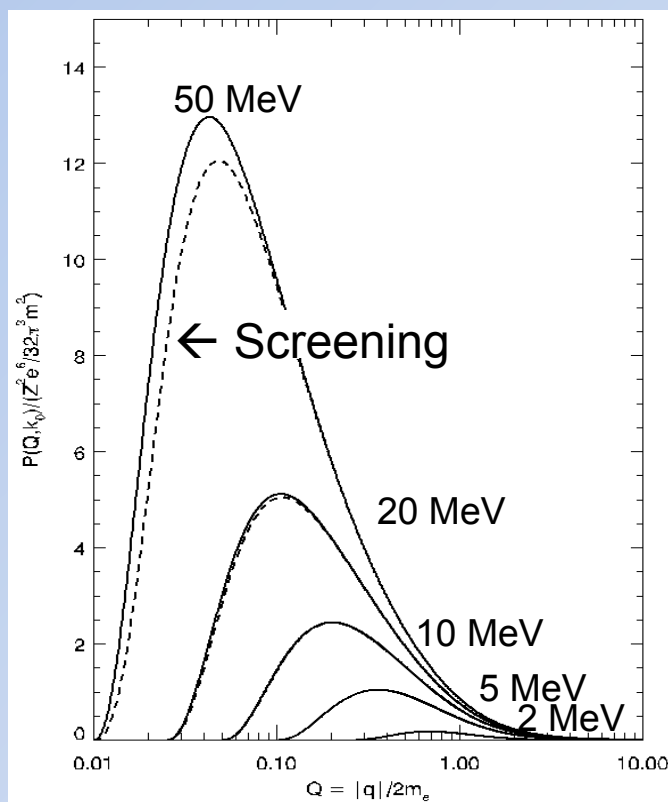
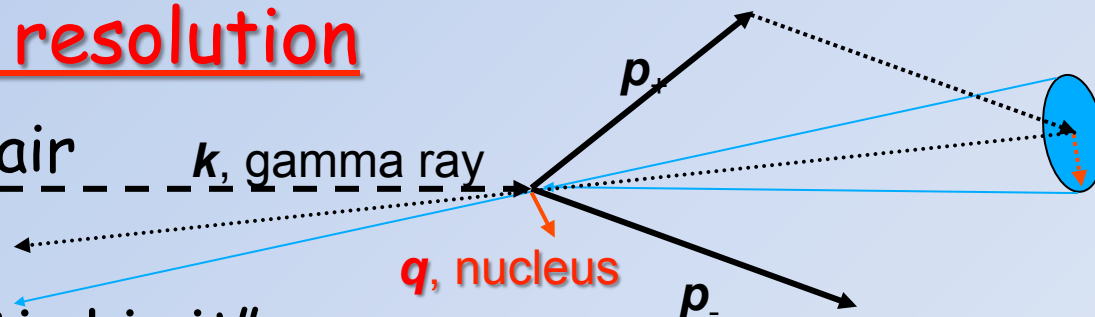
AdEPT Science, 5-200 MeV

- AdEPT will reveal the configuration of the most energetic accelerators in the Universe
- Explore fundamental processes of particle acceleration in active astrophysical objects
 - Pulsars, pulsar nebulae, supernova remnants, active galactic nuclei, magnetars, accreting binaries, gamma-ray bursts, ...
- Map the transition from electronic to hadronic processes in the Galactic diffuse emission
- Probe the universe for exotic processes

AdEPT Design Philosophy

- Optimize for angular resolution

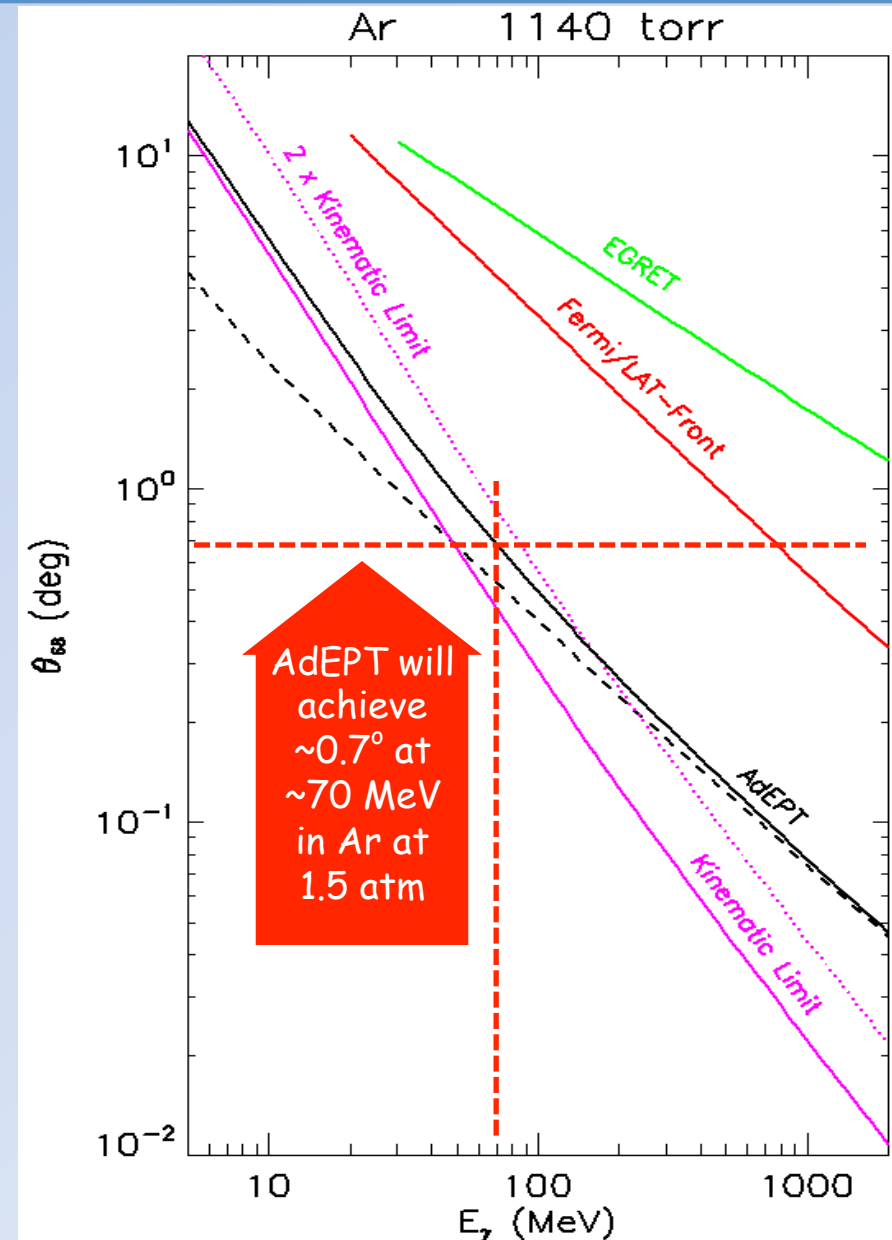
- Angular resolution of pair telescope limited by nuclear recoil, "Kinematic Limit"



AdEPT Angular Resolution

Achieving angular resolution near the Kinematic Limit
Continuous medium track imager density must be $< \sim 5 \text{ mg/cm}^3$
i.e. a Gaseous medium

Hunter et al., Astroparticle Physics 59, 18-28 (2014)



AdEPT Design Philosophy II

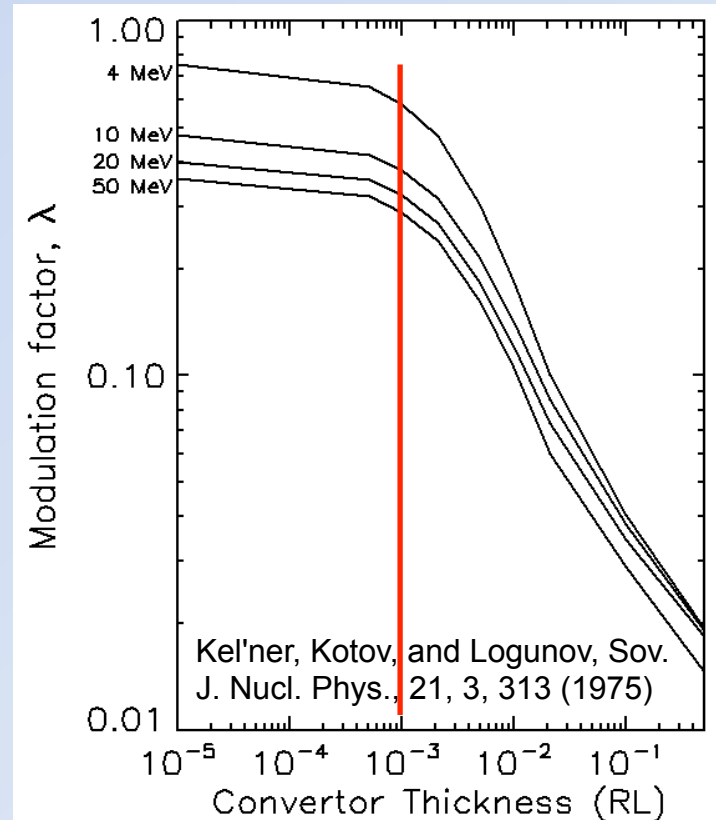
- Optimizing for angular resolution
→ polarization sensitivity

– Modulation factor, λ ,

$$\sigma(\varphi) = \frac{\sigma_0}{2\pi} [1 + P\lambda \cos^2(\varphi)]$$

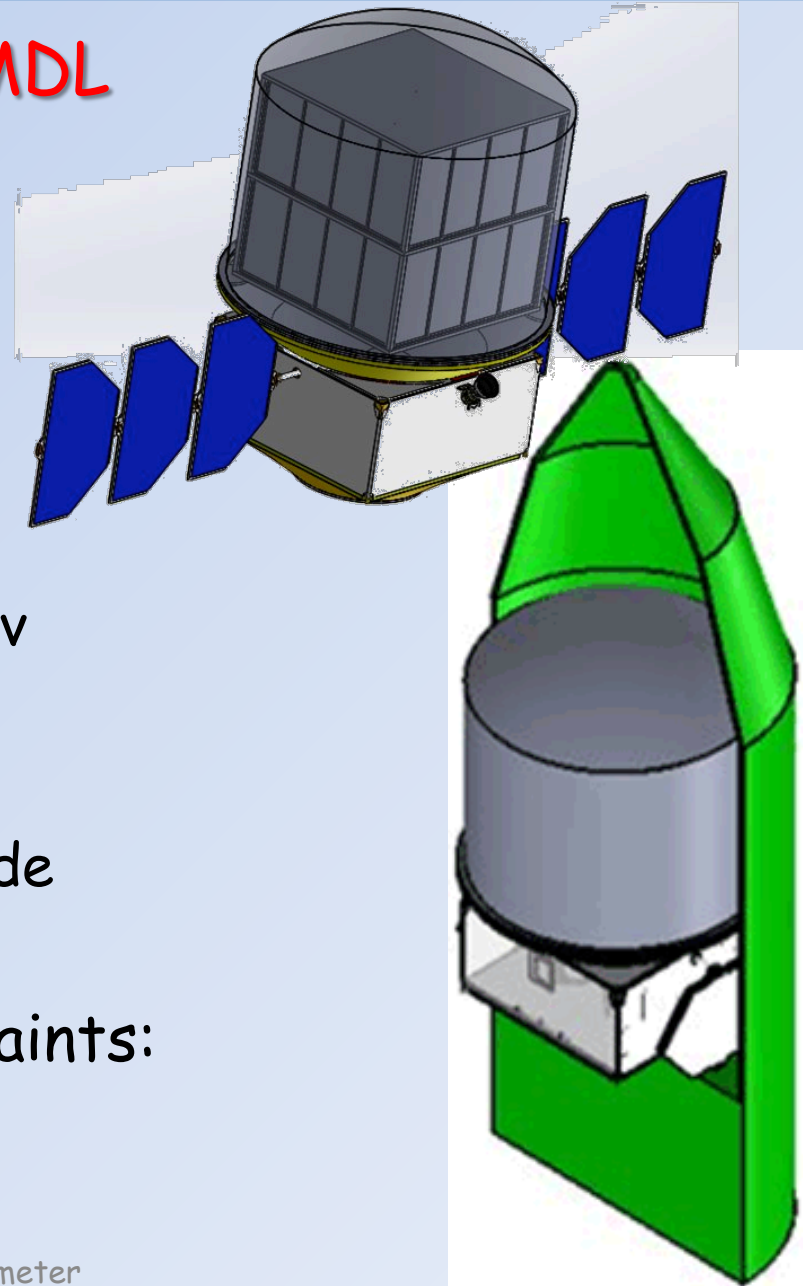
decreases exponentially with
thickness of tracking medium
above ~ 1 mRL

- Measure e- and e+ directions in $\lesssim 1$ mRL
 - $\sim 100 \mu\text{m}$ of Si
 - ~ 8 cm of Ar at 1.5 atm



AdEPT is a Viable Gaseous G-ray Polarimeter!

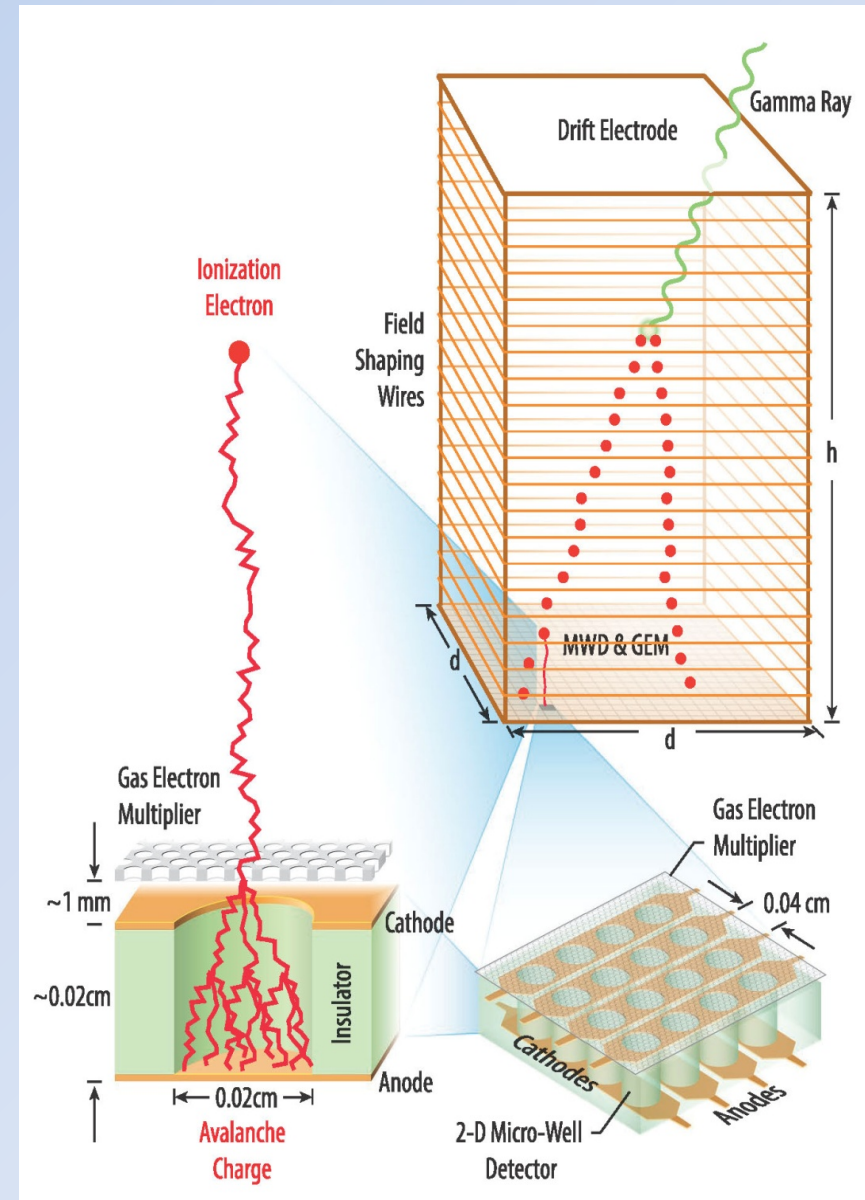
- **Baseline design studied in IDL/MDL**
- 2×2×2 array of 1 m³ 3-DTI modules
 - A_{geom} : 4 m², ~40,000 channels
- Ar (1100 torr) + CS₂ (40 torr), 25° C
- Pressure vessel: Al, 4 mm thick, ~300 cm diameter, ~530 kg
- Instrument power: ~500 W,
mass: ~320 kg w/o s/c, pv
- Spacecraft: zenith pointed, 3-axis stabilized, scanning mode
- Orbit: near equatorial, ~550 km altitude
- Athena launch vehicle
- Fits within Explorer mission constraints:
Mass, Power, & Cost



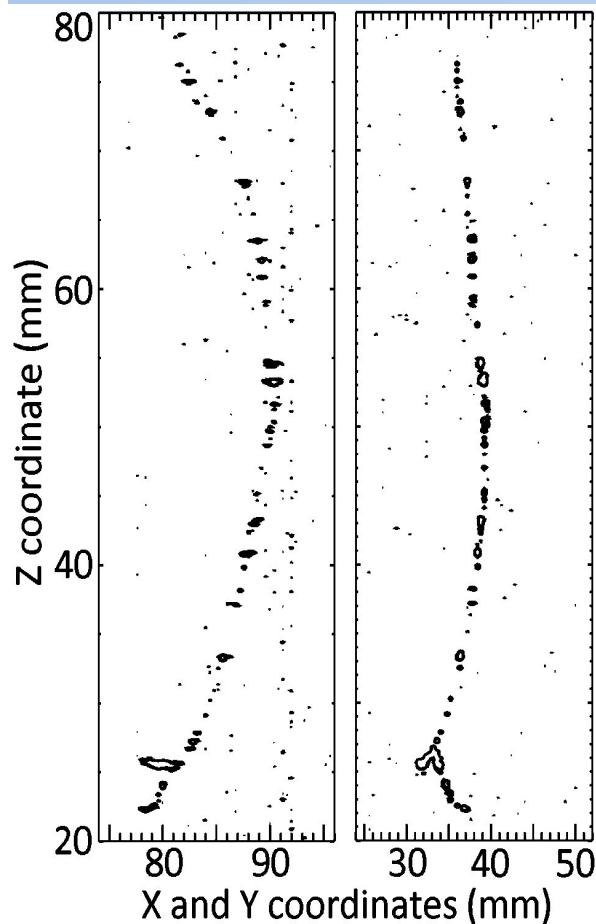
3-Dimensional Track Imager (3-DTI)

- **AdEPT Enabling Technology**

- Large-volume gas **time projection chamber (TPC)**
 - Low density, homogenous, 100% active particle tracking
 - Thermal diffusion achieved with negative ion drift
- **2-D readout, 2-D micro-well detector (MWD) + GEM**
 - Active detector, 0.4 mm pitch
 - GEM provides additional gain lost to negative ion drift
- Isotropic medium, scalable to large sensitive volume

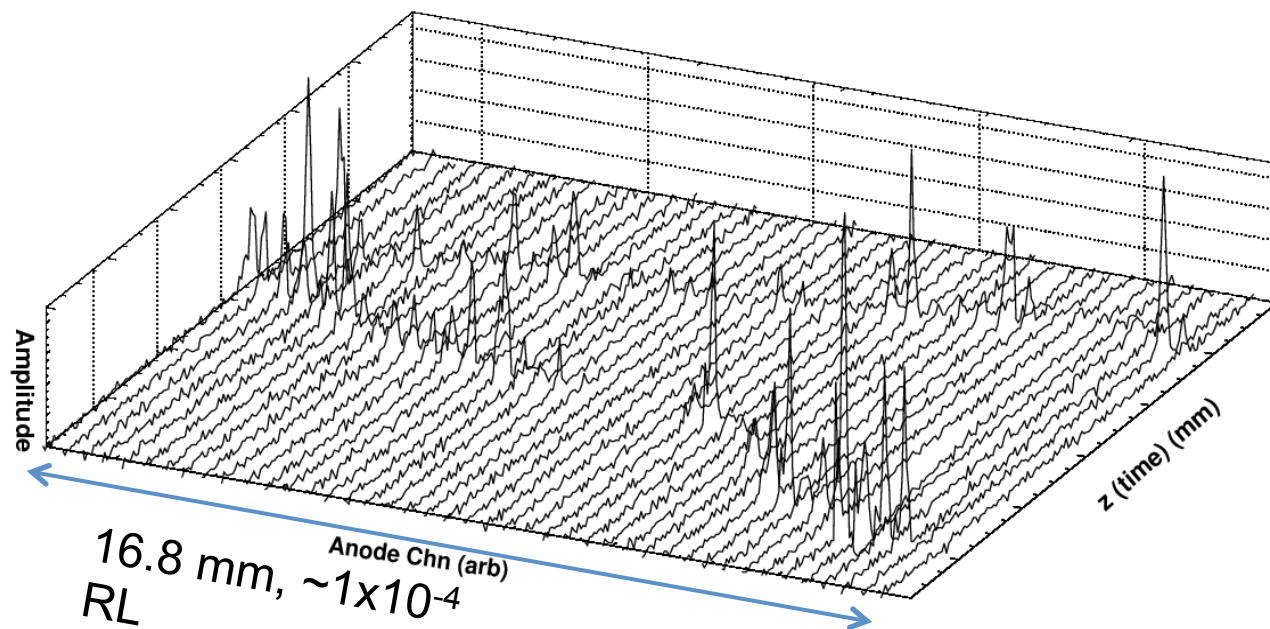


Electron Tracking in 3-DTI



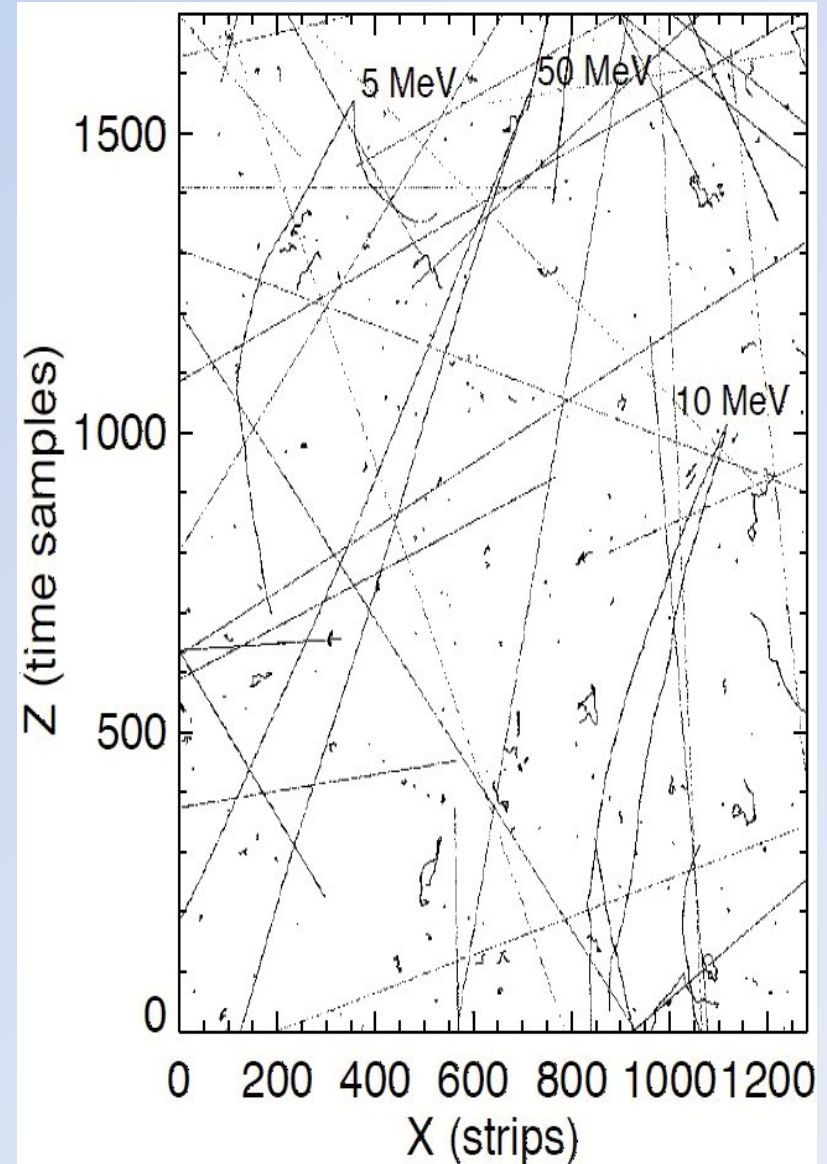
X-Z, & Y-Z projections of single electrons from ^{90}Sr in $\text{Ar} + \text{CS}_2$ with 0.4 mm resolution

X-Z projection of 6.129 MeV gamma interaction in 80% P-10 + 20% CS_2

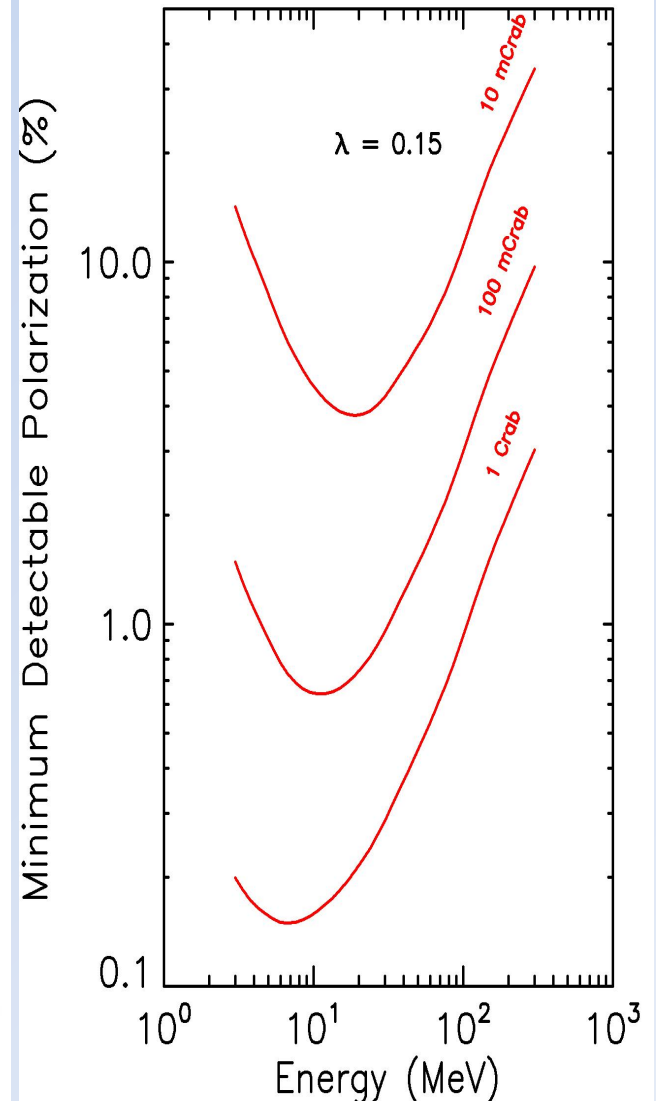
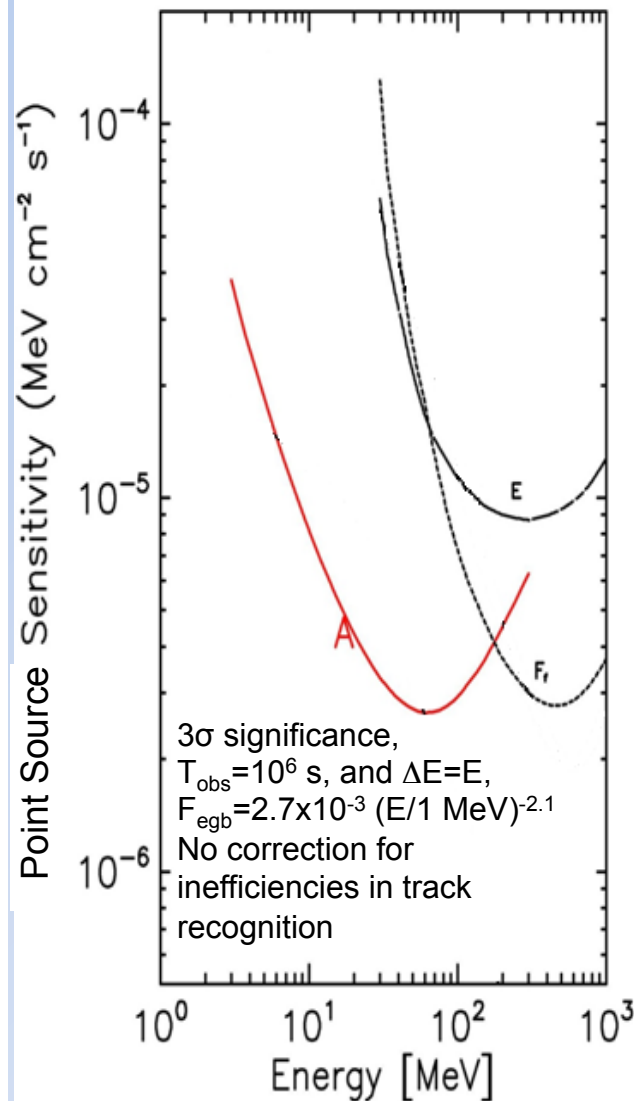
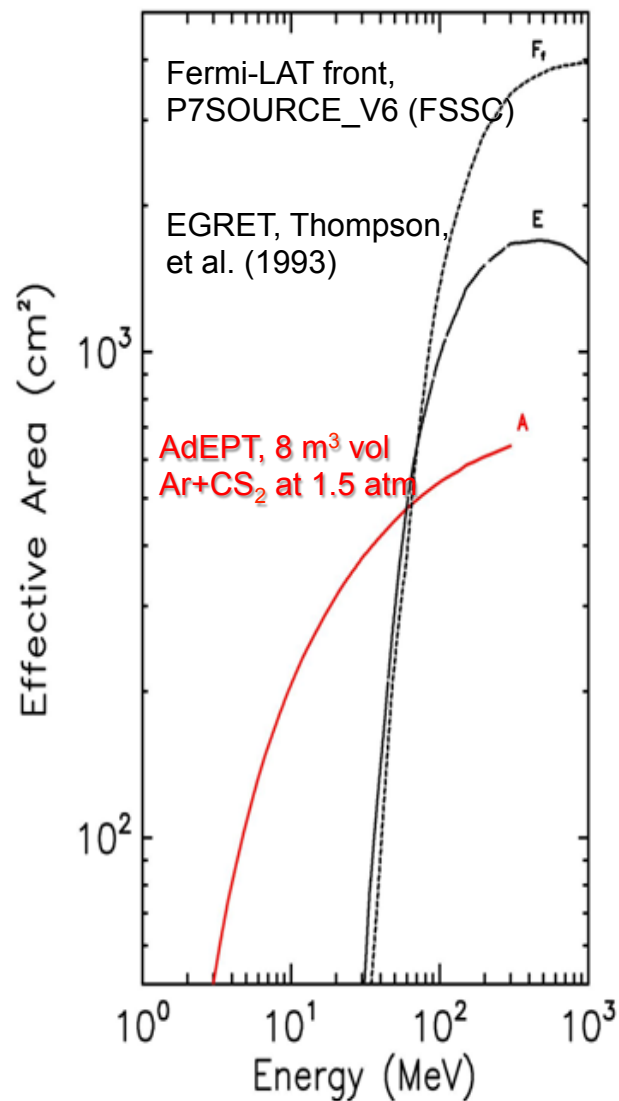


TPC Memory and Data Processing

- Ionization charge from all charged particles traversing TPC volume drifts to MWD and is "read out"
- Total drift time is ~ 50 ms
- Expect ~ 175 CR protons per m^3 in TPC volume
- Gb/s raw data rate
- Developing simulated data and multi-core on-board processing to discriminate gamma-rays from CRs, low-energy gamma interactions, & noise



AdEPT Baseline Performance



AdEPT Instrument Development

- 2015-18 ROSES-APRA
 - 50 x 50 x 100 cm³ AdEPT prototype
 - Multi-core processor to discriminate gamma-rays from background
 - Determine gamma-ray direction, energy, polarization, and time of arrival
 - Large area MWD integration
 - FEE ASIC
 - Calibrate at accelerator with polarized gamma rays, 5 - ~90 MeV
 - Determine electron energy from Coulomb scattering
 - Measure angular resolution and Polarization sensitivity
- Future NASA mission!

